

# Phyx 320

# Modern Physics

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February 8, 2021

Reading: 37.1 – 37.6

Homework #3 and Reading Reflection Due Tuesday 11:59 pm

# Relativity

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Asserted that all laws of physics are independent of reference frame

- Implies speed of light is constant

Required us to change our understanding of space and time to be frame dependent

Following this through made us change our definitions of momentum and energy

Makes it impossible to accelerate any object faster than the speed of light

Concluded that there's an energy associated with mass

$$x' = \gamma(x - vt)$$
$$t' = \gamma\left(t - \frac{v}{c^2}x\right)$$

$$p = \gamma_p m u$$

$$E = mc^2$$

# Early Evidence for QM

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Around the same time that Relativity was being developed, many observed phenomena could not be explained

Experiments into the nature of atoms did not follow electromagnetism or Newtonian mechanics

Needed a new description of atoms = Quantum Mechanics

# Interaction of Light and Matter

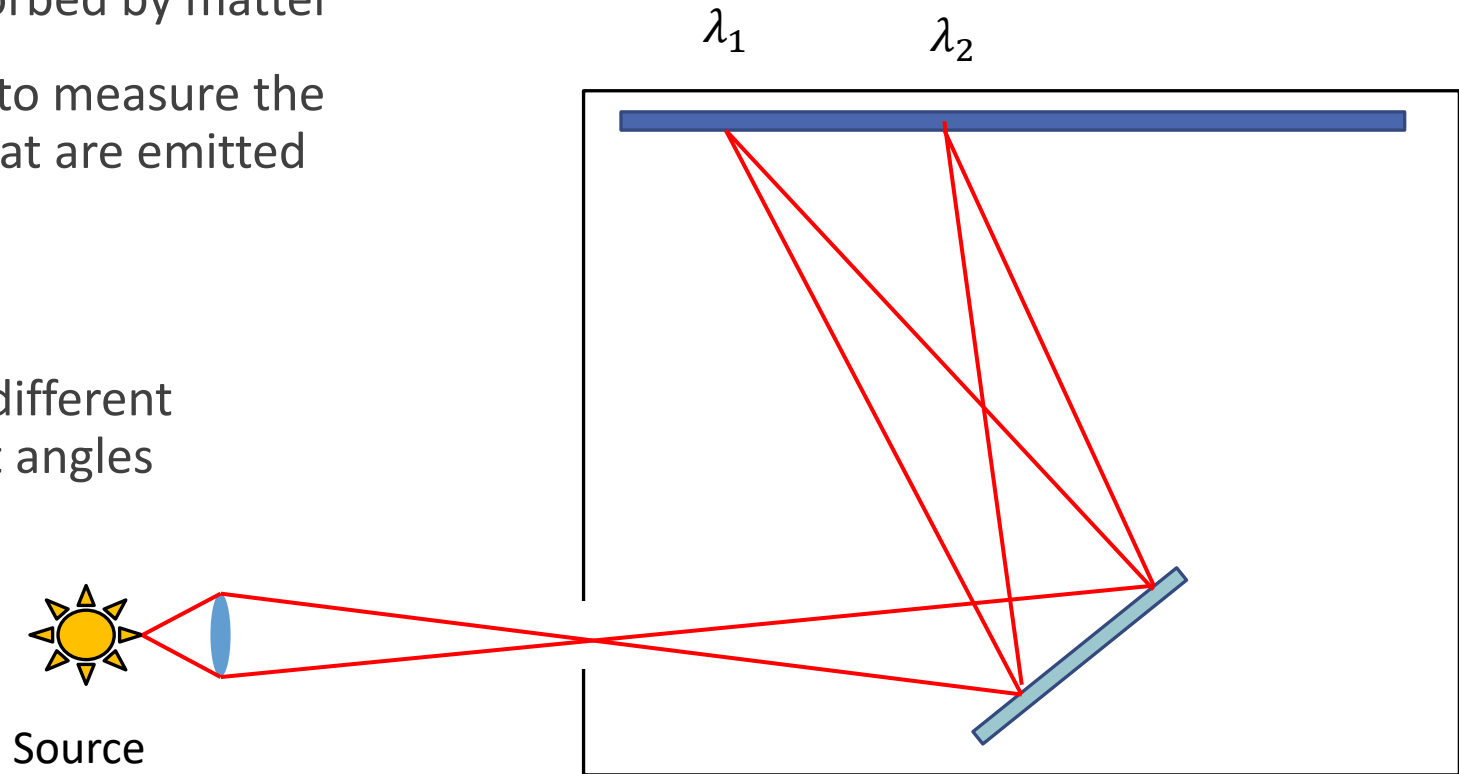
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Light can be emitted or absorbed by matter

We can use a spectrometer to measure the collection of wavelengths that are emitted (spectrum)

Wavelength = color of light

Diffraction grating diffracts different wavelength light at different angles



# Continuous Emission

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Hot objects emit light with a continuous spectrum

- All frequencies emitted at some level

This is how incandescent light bulbs work and why electric heating elements turn red

A spectrometer splits the light into a smoothly varying spectrum

Our eyes see the “average” of the spectrum



$\lambda = 400 \text{ nm}$

$\lambda = 700 \text{ nm}$

$$c = \lambda f$$

Wavelength

Speed of light

Frequency

The diagram shows the equation  $c = \lambda f$  with three blue arrows pointing to the variables: one from 'Wavelength' to  $\lambda$ , one from 'Speed of light' to  $c$ , and one from 'Frequency' to  $f$ .

# Blackbody Radiation

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Thermal radiation depends on

- Emissivity
- Surface area
- Temperature

Emissivity is how efficiently the object emits and absorbs

A blackbody is an object that absorbs light perfectly ( $e=1$ ) implies that it also emits light perfectly

The diagram shows the Stefan-Boltzmann law equation:  $\frac{Q}{\Delta t} = e \sigma A T^4$ . Blue arrows point from text labels to the corresponding variables in the equation: 'Heat energy per time' points to  $\frac{Q}{\Delta t}$ , 'Stefan-Boltzmann constant' points to  $\sigma$ , 'Emissivity' points to  $e$ , 'Surface Area' points to  $A$ , and 'Temperature' points to  $T^4$ .

Heat energy per time

Stefan-Boltzmann constant

$$\frac{Q}{\Delta t} = e \sigma A T^4$$

Emissivity

Surface Area

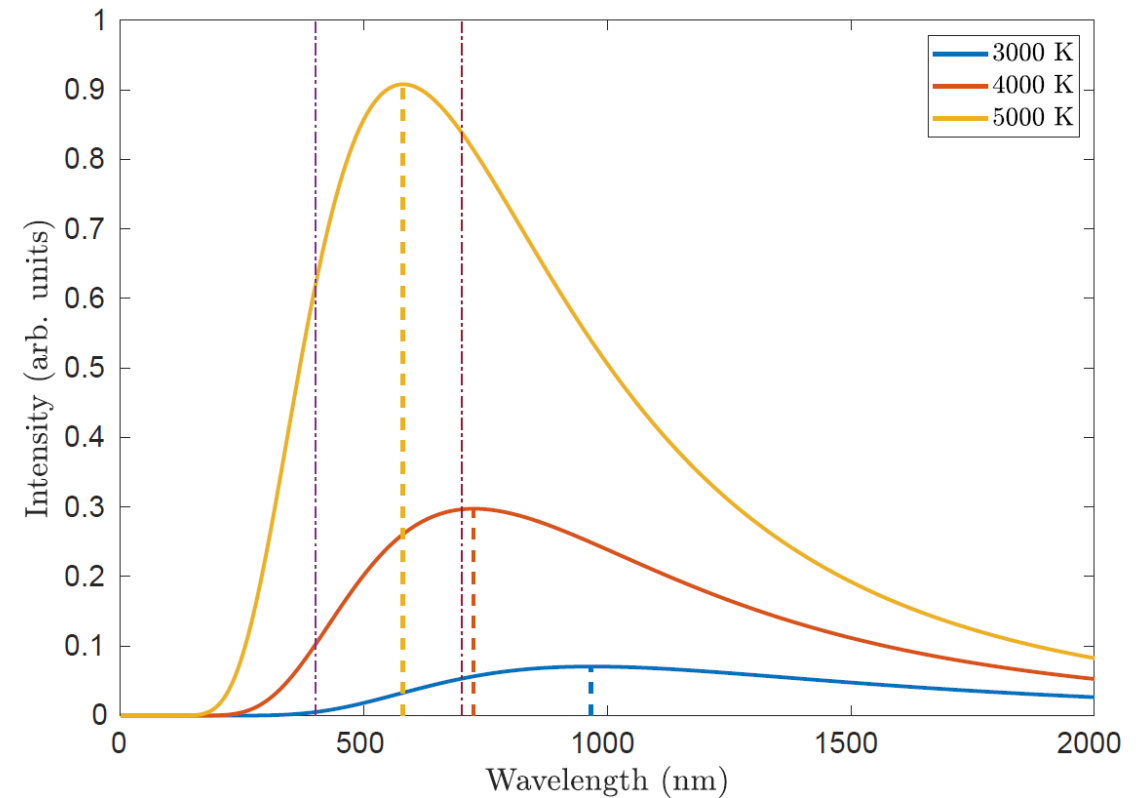
Temperature

# Blackbody Radiation

Blackbody (thermal) radiation is at all wavelengths (ultraviolet, visible, infrared)

- Spectrum only depends on temperature of the object
- Hotter objects emit more radiation at all wavelengths
- Higher the temperature the shorter the peak wavelength

No classical explanation



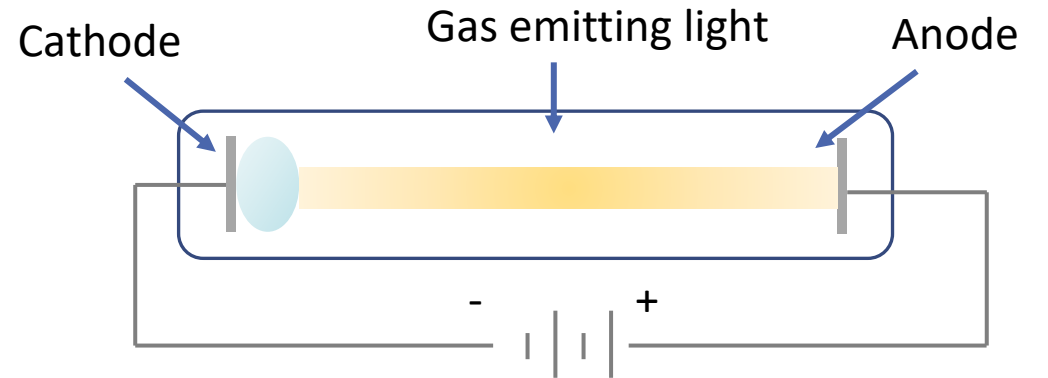
$$\text{Wein's Law: } \lambda_{peak} = \frac{2.90 \times 10^6 \text{ nm K}}{T}$$

# Discrete Spectra

If we put gas in discharge tube, we get a discrete spectrum

- Depends on what element is in the tube
- This is how neon lights and florescent lighting works

Looking at emission spectrum we can know which element is in the tube (elemental fingerprint)



Emission Spectrum





# Discrete Spectra

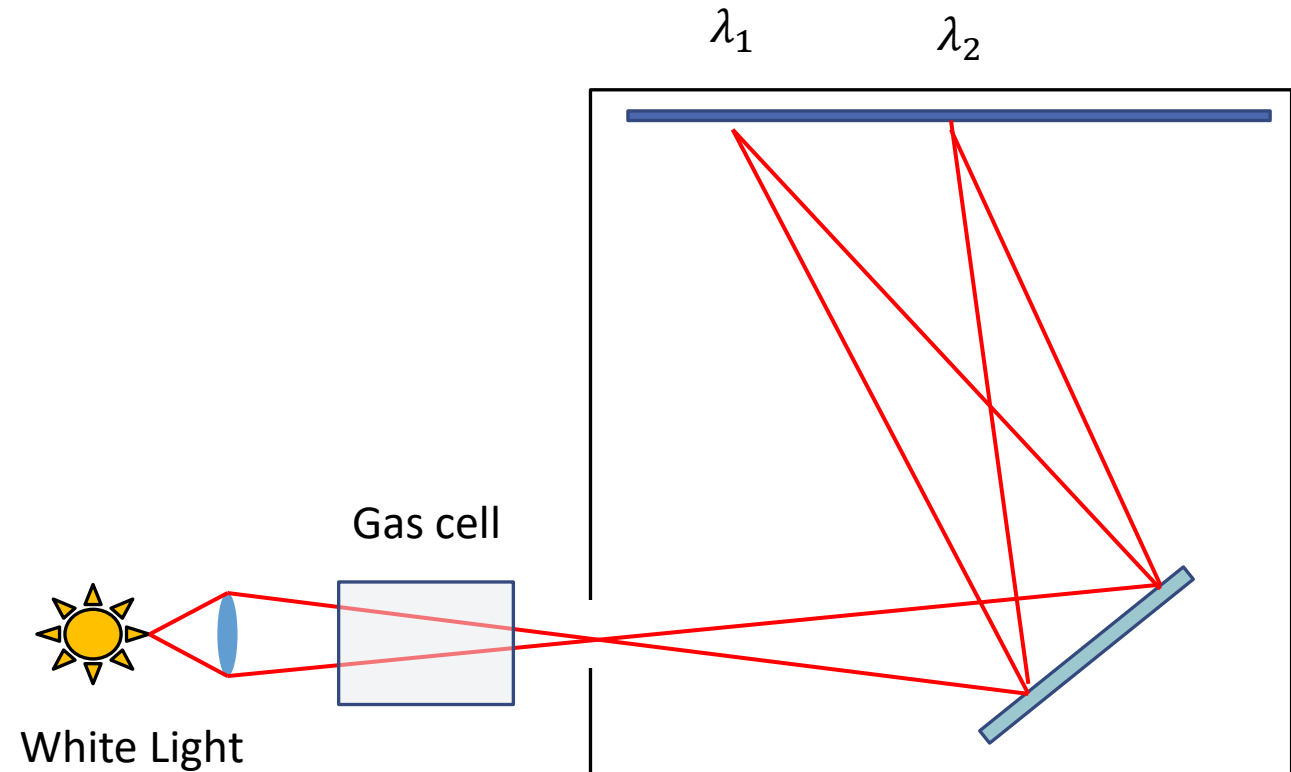
If we send white light through gas, we also get a discrete spectrum

- Depends on what element is in the tube

Every wavelength that is absorbed is also emitted but not every emitted wavelength is absorbed

**What causes atoms of different element absorb and emit different, discrete spectra?**

Absorption Spectrum



# Balmer Series

When looking at hydrogen Balmer realized that the spectrum could be described by an empirical law:

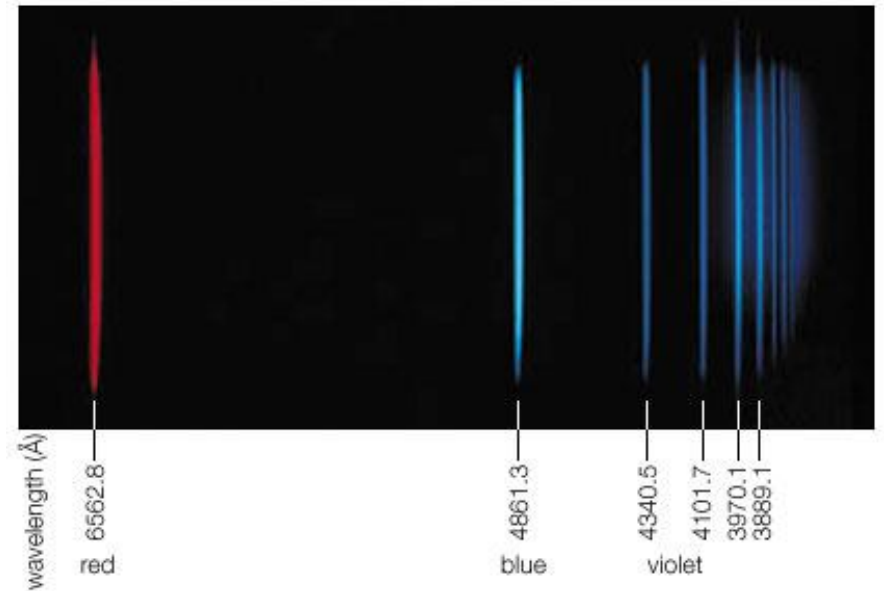
$$\lambda = \frac{91.18 \text{ nm}}{\frac{1}{2^2} - \frac{1}{n^2}}$$

$$n = 3, 4, 5, \dots$$



Discrete Series

Emission spectrum of Hydrogen



# Balmer Series

With more sophisticated measurements, more groups of emission lines were found.

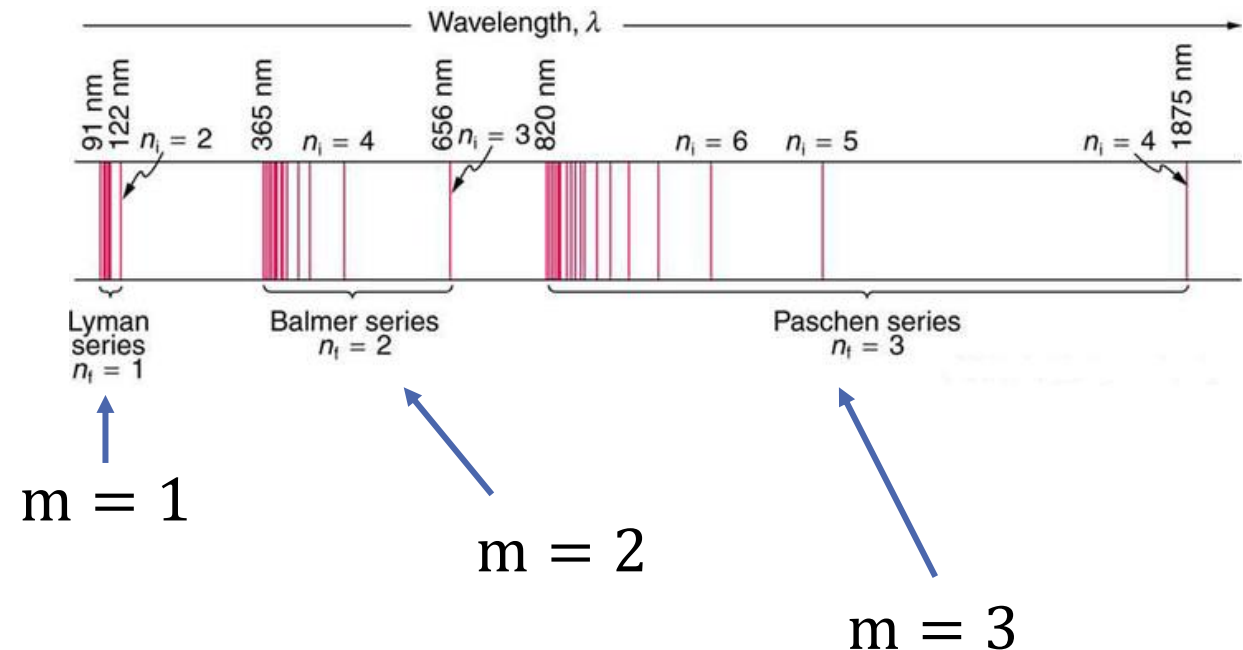
Balmer's result was generalized to:

$$\lambda = \frac{91.18 \text{ nm}}{\frac{1}{m^2} - \frac{1}{n^2}}$$

$$m = 1, 2, 3, \dots$$

$$n = m + 1, m + 2, \dots$$

Emission spectrum of Hydrogen



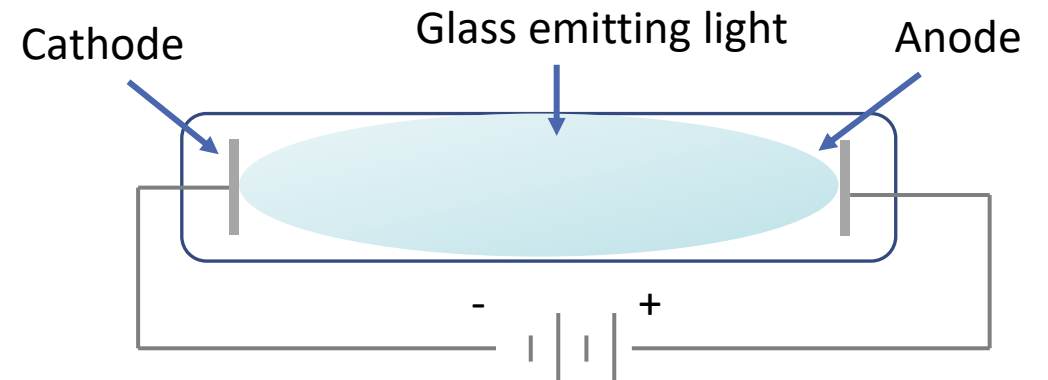
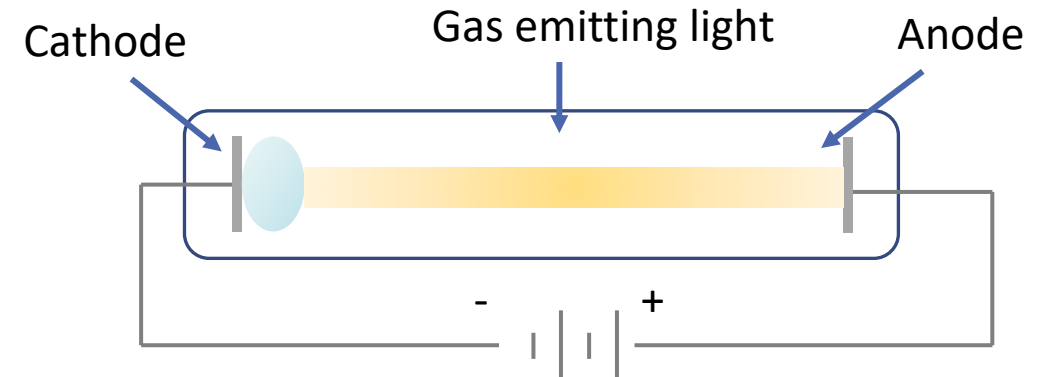
# Cathode Rays

As vacuum technology got better, experimenters found that a discharge tube with no gas (vacuum) still emits light

Some sort of rays were emitted by the cathode (named cathode rays)

When these rays hit the glass, they caused it to emit green light

Used in old TVs and computer monitors



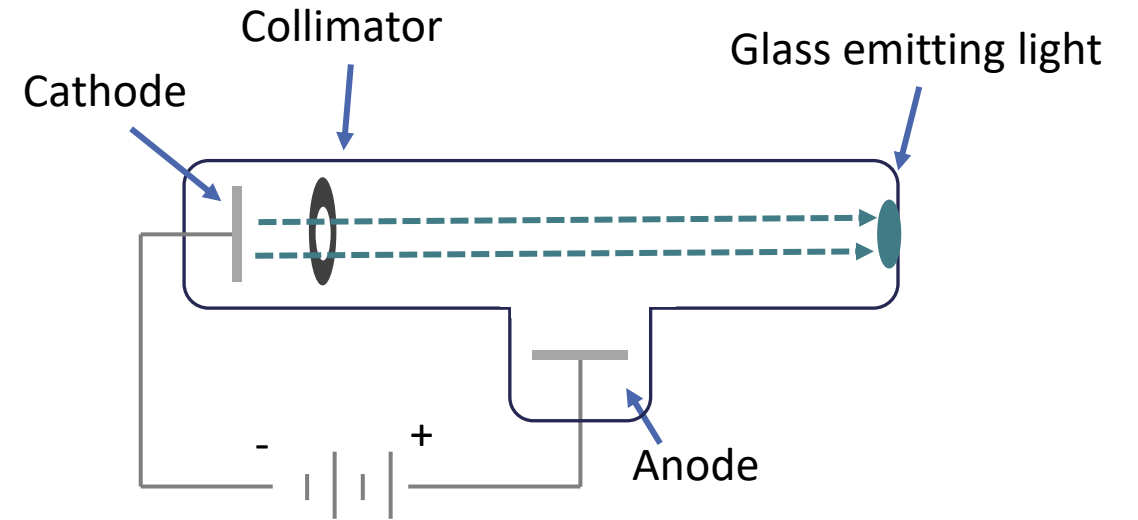
# Crookes Tube

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Crookes added a collimator to form a beam of cathode rays

Found that the rays did not depend on cathode material

Early version of a particle accelerator



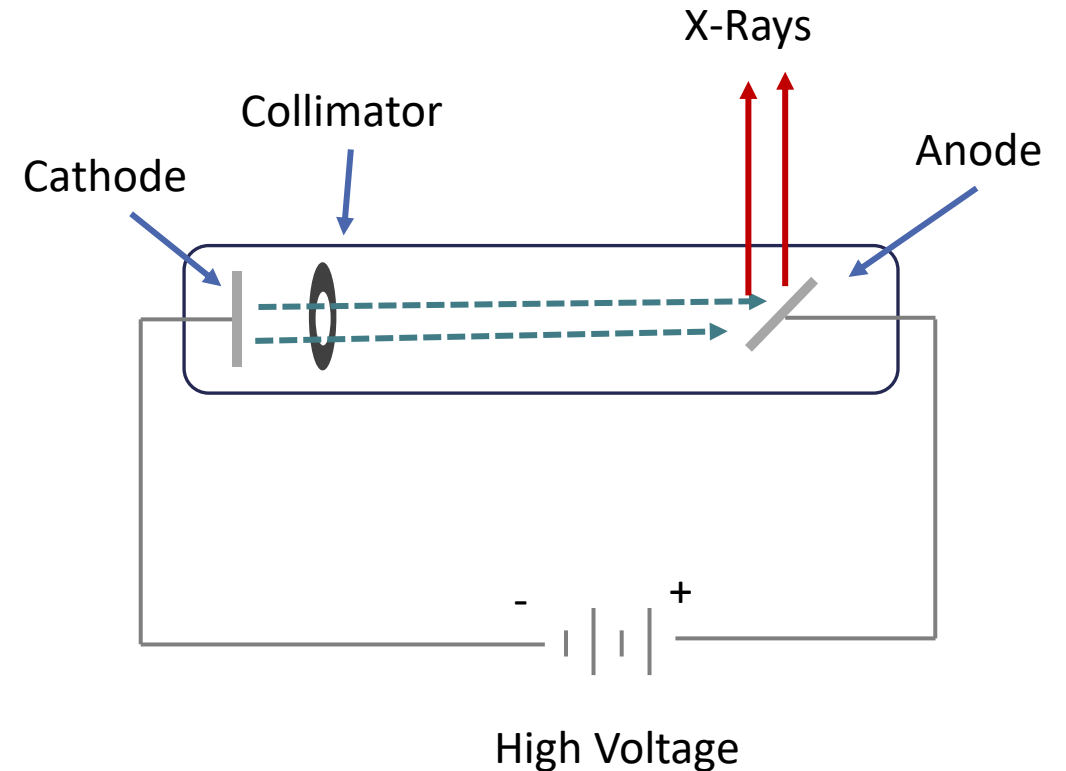
# X-Rays

Röntgen decided to increase the voltage and let the cathode rays strike the anode in vacuum

Produced x-rays for the first time

Found that they pass through most material and traveled in straight lines

20 years later, it was shown that x-rays are EM waves with wavelengths of 0.01 nm – 10 nm



# Homework Questions

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